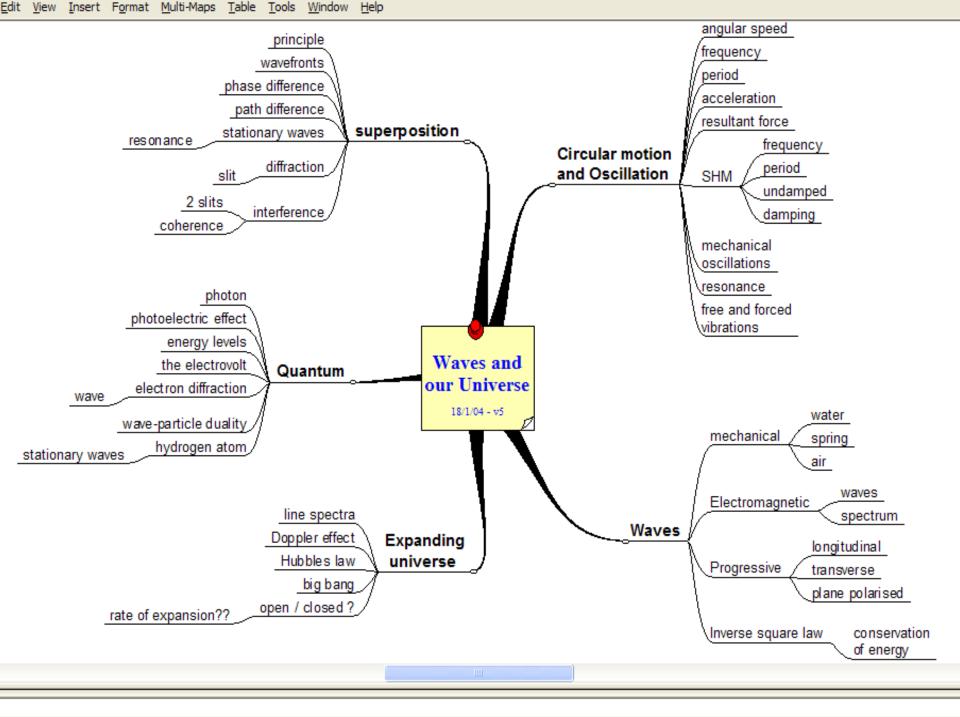
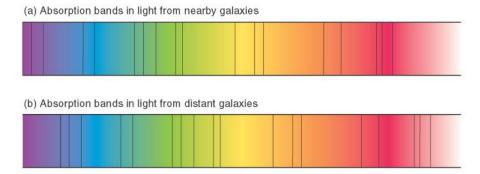
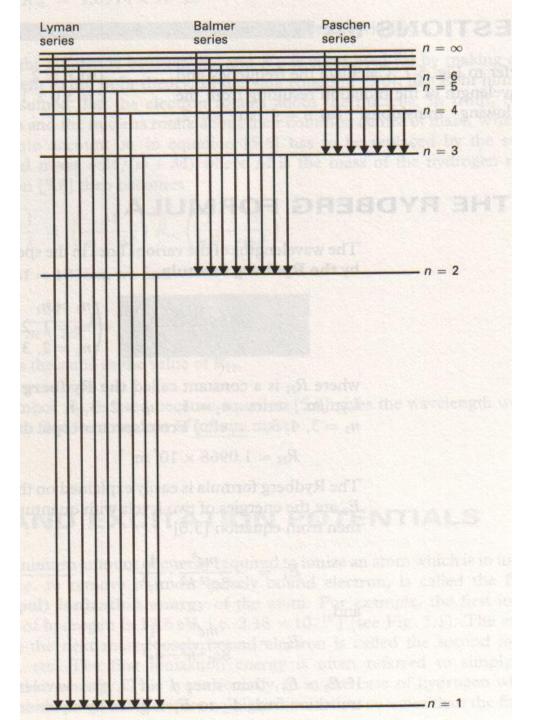
Our Universe

Dr P Lawson



Line Spectra





mooning ugin.

SOURCES OF LIGHT

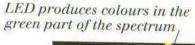


BRIGHT FILAMENT LAMP

This spectrum shows which colours are produced



All colours of light together combine to produce white BRIGHT FILAMENT LAMP With a high electric current, the whole spectrum of visible light is

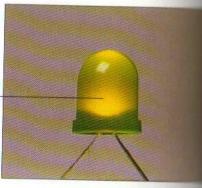




LED appears green

GREEN LED

An LED (light-emitting diode) is made of a semiconductor, and produces certain colours of light.



GREEN LED

DIM FILAMENT LAMP

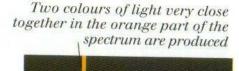
Red, yellow, and green light combine to produce orange



Lamp appears orange

No blue light produced

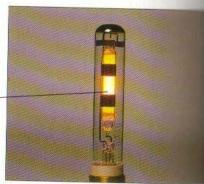
DIM FILAMENT LAMP With a smaller current, the temperature of the filament (see pp. 32-33) is low.



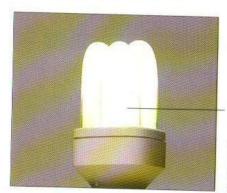
Lamp appears orange

SODIUM LAMP

In a sodium lamp, an electric current excites electrons in sodium vapour, giving them extra energy. The electrons give the energy out as light.



SODIUM LAMP



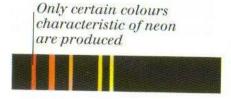
FLUORESCENT LAMP

Lamp produces certain colours in each part of the spectrum



All three types of cone are stimulated and lamp appears white

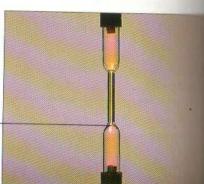
FLUORESCENT LAMP In a fluorescent lamp, chemicals called phosphors produce colours in many parts of the spectrum.



Lamp appears orange

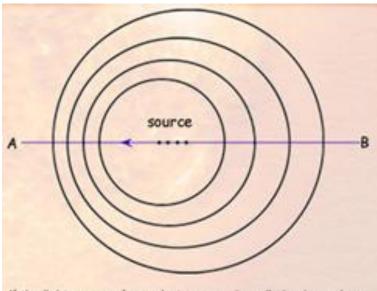
NEON TUBE

In a similar way to a sodium lamp, a neon discharge lamp produces a characteristic orange glow.



NEON TUBE

Doppler effect



If the light source of any electromagnetic radiation is moving toward point A, the wavelengths are compressed (blue shifted) whilst wavelengths from source to B are extended (red shifted).

..\..\..\Desktop\CDs\physics applets\ph11e\dopplereff.htm

By comparing the absorption lines produced in a star's or galaxy's spectrum by a variety of chemical elements, with laboratory produced line spectra for the same elements, a line shift and so wavelength shift can be measured.

When a light source is moving away from the Earth, its wavelength becomes stretched. The increase in wavelength produces a red shift in the spectrum. In contrast, when a light source is moving toward the Earth, its wavelength becomes compressed. The decrease in the wavelength produces a blue shift in the spectrum.

Doppler Effect

The greater the velocity of the light source relative to the Earth, the greater is the amount of red or blue shift observed. For all electromagnetic radiation this Doppler effect is given by,

$$\frac{\Delta f}{f} = \frac{\Delta \lambda}{\lambda} = \frac{\nu}{c}$$
 when the velocity of the light source, ν is much less than c

where Δf and $\Delta \lambda$ are the changes in frequency and wavelength, and f and λ are the frequency and wavelength of the emitted electromagnetic radiation.

For light which is red shifted, as the wavelength increases, its frequency must decrease. For light which is blue shifted, as the wavelength decreases, its frequency must increase.

Example

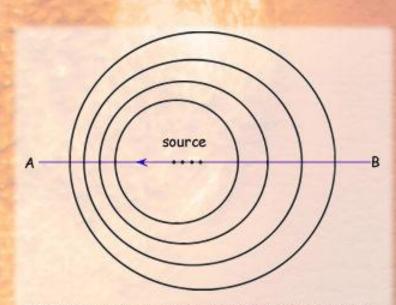
If the hydrogen absorption line in the spectra of a star is 6.6570×10^{-7} m, and the same element in a laboratory test is 6.5628×10^{-7} m then the velocity of the star can be found from the Doppler effect.

$$\frac{\Delta \lambda}{\lambda} = \frac{v}{c} = \frac{(6.5628 - 6.6570) \times 10^{-7}}{6.5628 \times 10^{-7}} = 0.01435$$

$$v = -0.01435c$$

$$v = -4.3 \times 10^{6} \text{ms}^{-1}$$

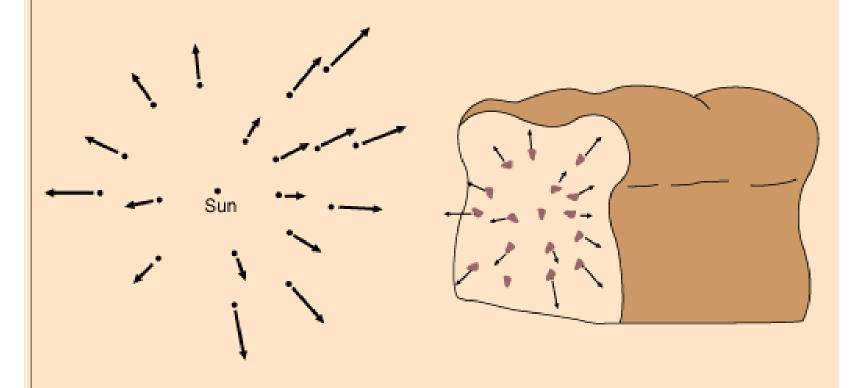
The wavelength in the star's spectra is longer (red shifted) so the star's radial speed must be away from the Earth.



If the light source of any electromagnetic radiation is moving toward point A, the wavelengths are compressed (blue shifted) whilst wavelengths from source to B are extended (red shifted).

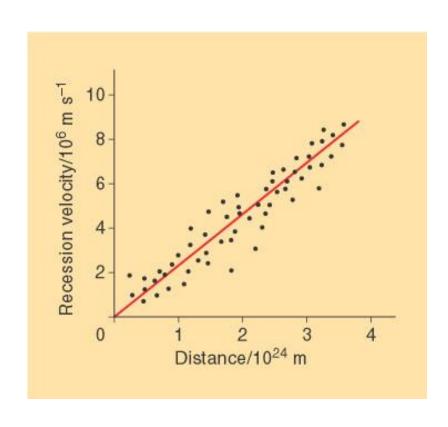
Expanding Universe

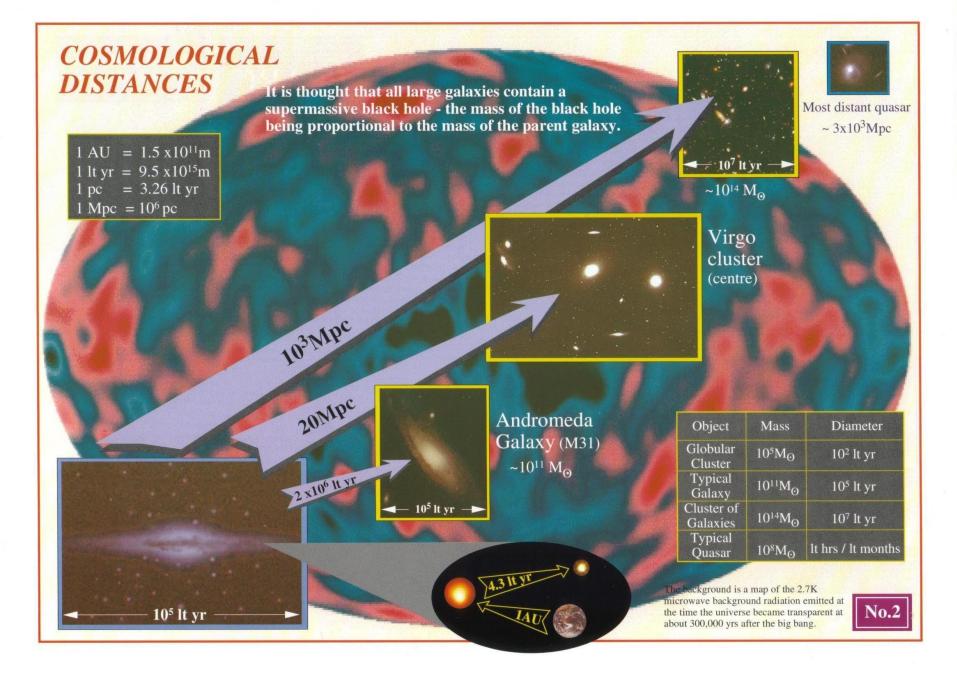
The galaxies we see in all directions are moving away from the Earth, as evidenced by their <u>red shifts</u>. <u>Hubble's law</u> describes this expansion.

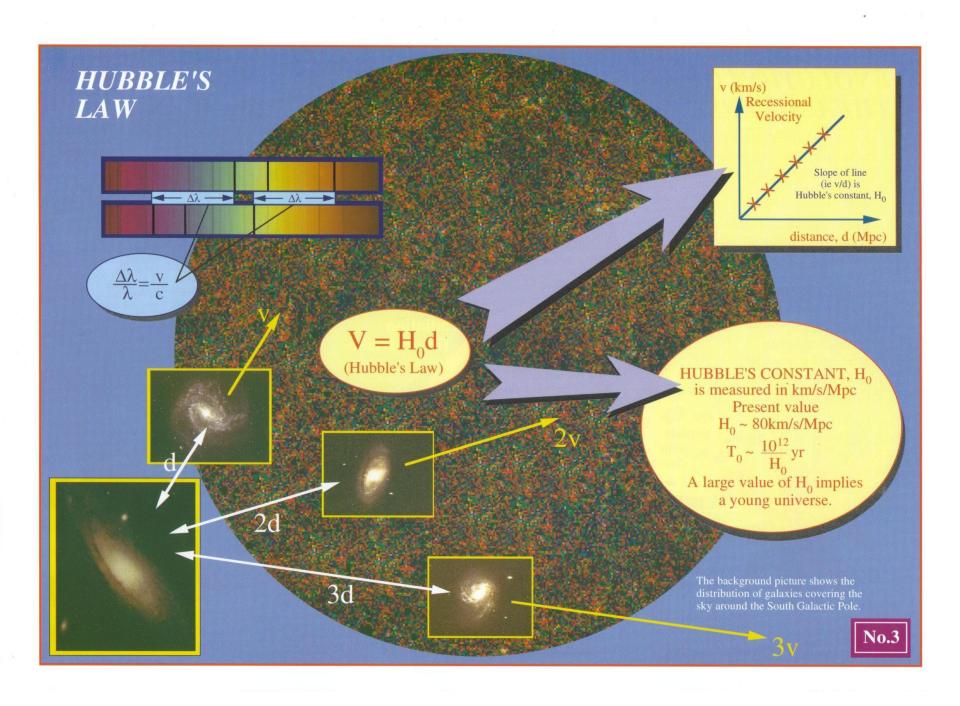


The fact that we see all stars moving away from us does <u>not</u> imply that we are the center of the universe! <u>All</u> stars will see all other stars moving away from them in an expanding universe. A rising loaf of raisin bread is a good visual model: each raisin will see all other raisins moving away from it as the loaf expands.

Hubbles Law

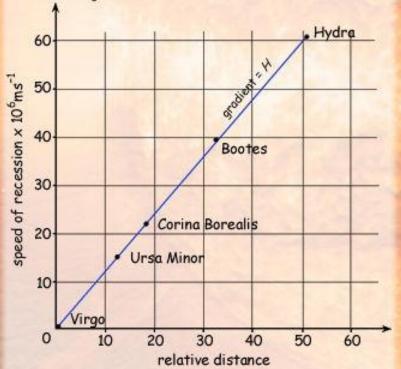






Galaxies are observed to have line spectra, that are red shifted. They are therefore considered to be receding from the Earth, and from each other. The only exceptions to this are the local galaxies, some of which have blue shifts. This is because of their rotational motion relative to each other which blurs any recessional motion.

The graph shows the increase in recessional speed with increased distance for the five brightest clusters in each of the galaxies named.



The constant H, probably lies in the range 50 - 100kms-1Mpc-1

Hubble's law states that the speed u, of a galaxy's recession is proportional to its distance, given by,

v = Hd where H = Hubble's constant, and d = distance from Earth.

This means that the more distant a galaxy is, the greater is its recessional velocity and so the greater is its red shift.

Hubble's Law

time =
$$\frac{\text{distance}}{\text{velocity}} = \frac{d}{v}$$
 therefore, $t = \frac{1}{H}$

As time began with the Big Bang when space began to expand, then the age of the universe is:

H

This assumes that the Hubble constant H, has remained constant through time. For this to be true the universe would have had to expand at a constant rate. The actual rate may be slowing due to gravitational effects. If, however, a value of 50 kms⁻¹Mpc⁻¹ is taken, then that means a galaxy 1Mpc away is receding at 50kms⁻¹.

As
$$1pc = 3.086 \times 10^{16} m$$
, therefore $1Mpc = 3.086 \times 10^{22} m$

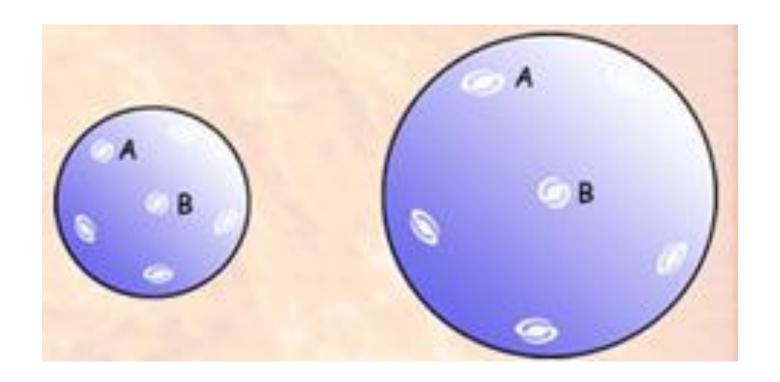
This means that a galaxy 3.086 × 10²²m away is receding at 50 kms⁻¹. If this has always been the rate then recession has occurred for,

$$t = \frac{3.086 \times 10^{22}}{50 \times 10^3} = 6.0 \times 10^{17} \text{s}^{-1}$$
 This is about 20 billion years.

The value of Hwhich is often taken is H=75 kms⁻¹Mpc⁻¹.

This gives an age for the universe of 15 billion years.

Big Bang





The diagram represents the changes in the universe from the Big Bang to the present day.

Upper section shows the hot dense clumps of matter that became galaxies.

10-10_S

The lower section shows radiation and subatomic particles that later became atoms and eventually plants and animals living on Earth.

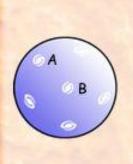
15 billion years life on earth, 1 billion years heavy 1 million years elements formed stars and galaxies in stars 300,000 years microwave background radiation 3 minutes fills universe helium nuclei forme 1 second neutrons and protons $10^{15} deg$ $10^{10} deg$ 109deg This series of 4 summary sheets has been produced at Royal Holloway, University of London with financial support from the Particle Physics and Astronomy Research Council (PPARC). They may be freely copied or additional copies can be obtained from: -270 The Physics Secretary, Physics Department, Royal Holloway, University of London, Egham, Surrey, TW20 0EX Tel: 01784 443448 Copies are also available from our department web site at http://www.ph.rhbnc.ac.uk/ No.4

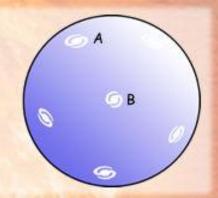
Further information about PPARC can be obtained from the Public Relations Office, PPARC, Polaris House, North Star Avenue, Swindon SN2 1SZ. Tel: 01793 442098

The Big Bang theory is an attempt to explain how the Universe came into existence, and what has happened to it since. According to the model if the Universe has been expanding for a time $t = \frac{1}{H}$, then it is this time ago when time and space must have started. Before this, space-time must have been infinitely curved. The Universe must have been incredibly small and had infinite density. Any event, which may have taken place before the Big Bang would be irrelevant to the physical laws we see in the Universe today.

If the value of $H = 75 \text{ kms}^{-1}\text{Mpc}^{-1}$, then space and time began 15 billion years ago at what is called the Big Bang singularity.

The expansion of space can be thought of like the expansion of the surface of a blown up balloon. There is no limiting boundary, and no centralised point.





The Big Bang

Evidence for the Big Bang theory include the following predicted features which have all been verified by observation:

- Background microwave radiation. This is received from all directions and is believed to be surviving black body radiation following the Big Bang which has cooled to 2.7K
- The ratio of hydrogen to helium is 3:1
- The red shifts of galaxies and their speeds of recession are proportional to distance, with the
 most distant galaxies (quasars) having the largest red shifts and highest recessional speeds.

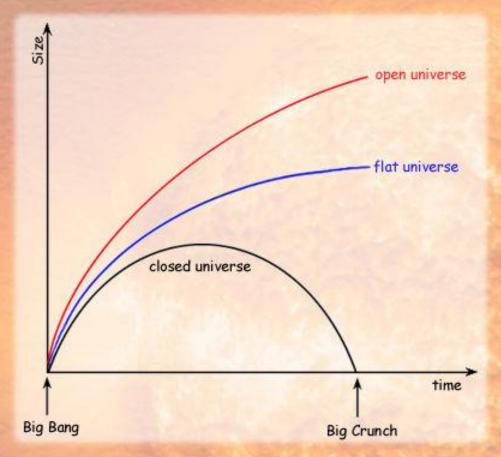
According to the Big Bang:

- Within the first millionth of a second fundamental particles formed. The universe inflated rapidly, and nuclear protons and neutrons formed from quarks. The temperature began to fall to ≈10¹³K
- Within a hundredth of a second matter-antimatter nucleon collisions had annihilated virtually all the less abundant antimatter nucleons
- Within a few seconds electron-positron collisions had annihilated most positrons
- Within a minute or two helium nuclei formed as some protons and neutrons joined as the temperature fell to ≈10⁹K.

At this point the universe would have been very small, very dense and not too dissimilar to the plasma state conditions found in the core of a star.

 Over the next few hundred thousand years the temperature fell to 3000K allowing atoms to form as electrons combined with nuclei. This allowed photons of em radiation to travel unhindered. (Background radiation detected today is thought to be the cooled remnants of this early photon energy). Stars and galaxies started to form under gravitational attraction.

Rate of expansion? Open or closed



As the Hubble constant H, is not known exactly, the value for the critical density is not known. From estimates made from visible matter, that is, matter that emits electromagnetic radiation, the average density of matter is well below the critical density. On this evidence the universe is likely to continue to expand as an open universe.

From the theory of relativity the relationship between critical density and the Hubble constant is given by, $p_c = \frac{3H^2}{8\pi G}$ A higher calculated value for the Hubble constant will increase the value for the critical density, and so make an open universe more likely.

At present the ultimate fate of the universe is not known. There are three possible outcomes, but which one actually takes place depends on how the average density ρ , of the universe compares with the critical density ρ .

The universe may be open, that is, it will continue to expand forever. This will occur if the average density of matter in the universe is less than the critical density $(p < p_c)$.

The universe may be flat, that is, the rate of expansion slows until the universe reaches a point where no more expansion takes place. This will occur if the average density of matter in the universe is equal to the critical density $(p = p_c)$.

The Universe may be closed, that is, it will reach a maximum expansion and then contract under the influence of gravity. This will occur if the average density of matter in the universe is greater than the critical density $(p > p_c)$.

Fate of the Universe

The actual amount of matter in the universe may, however, be much higher than the value found from visible matter. There may be a huge amount of dark matter not seen. Neutrinos may have a finite mass and be sufficiently abundant to allow the average density to equal or even surpass the critical density. The universe therefore could at some time in the future begin to contract. This would eventually lead to a Big Crunch.

EXPANSION

In a thousandth of the time it takes to blink an eye the universe expanded to the size of a grapefruit.

This expansion is equivalent to a molecule expanding to the size of a galaxy in the same time period.

New evidence suggests that the rate of expansion is **increasing**.

(This suggests a type of antigravity pushing things

time

Big Crunch

apart)

Dark Matter?
Neutrinos?
String theory?
M Theory?
Quantum Gravity?
Consciousness?